

Normalized water-leaving radiance (IOCCG-6 Meeting, 24-26 January 2001)
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The “Normalized water-leaving radiances ($nL_w(\lambda)$) are the products generally distributed by the various Space Agencies. They belong to the Level 2 products category.

Actually the radiances (and related “reflectances”), in spite of the normalization, are still depending on the pixel-sun geometry, and therefor are not straight forwardly comparable. This is due to the non-isotropic character of the upward radiance field.

A meaningful use of these $nL_w(\lambda)$, [such as the reproduction of Level 3, the intercomparison between sensors, the merging of data, and even the composition with sea-truth radiometric measurements] is not possible as long as the artifacts resulting from the bi-directional effects have not been removed.

Such a removal is easy in terms of computational effort (the correction is extracted from LUT). It is recommended that the various Agencies make available normalized water-leaving radiances, converted in the same way (same LUT), and in parallel with the classical product. The physics behind the corrections is well understood. The parameterization, however, may evolve according to progress in marine optics. This is the main reason for keeping the classical nL_w product, in such a way that any user can easily correct it according to his own scheme.

Basic equations

(simplified)

$$\begin{aligned}
 \checkmark \quad R &= \frac{E_u(0^-)}{E_d(0^-)} = f \frac{b_b}{a} \quad (= f' \text{ (IOP's)}) \\
 \checkmark \quad E_d(0^-) &= E_d(0^+) \frac{1 - \bar{\rho}}{1 - \bar{r}R} \\
 \checkmark \quad Q &= \frac{E_u(0^-)}{L_u(0^-)} \\
 \checkmark \quad L_w &= L_u(0^-) \frac{1 - \rho_F}{n^2} \\
 L_w &= E_d(0^+) \mathfrak{R} \frac{f}{Q} \frac{b_b}{a}
 \end{aligned}$$

where

$$\mathfrak{R} = \left[\frac{(1 - \bar{\rho})}{(1 - \bar{r}R)} \frac{(1 - \rho_F)}{n^2} \right] \quad (\text{reflection + refraction})$$

and

$$E_d(0^+) = \mu_0 \varepsilon_c F_0 t$$

t : atmospheric transmittance

F_0 : extra-terrestrial solar flux (mean Earth-Sun distance : d_0)

$\mu_0 = \cos(\theta_0)$, θ_0 = solar zenith angle

$\varepsilon_c = \text{eccentricity correction} = (d_0 / d)^2$

Expanded basic equations

- Re-introducing wavelength, geometry, and Chl (Case 1 waters)
- Wind speed and aerosol optical thickness dependencies ignored

$$\checkmark \quad R(\theta_0, \lambda, \text{Chl}) = \frac{E_u(0^-)(\theta_0, \lambda, \text{Chl})}{E_d(0^-)(\theta_0, \lambda, \text{Chl})^*} = f(\theta_0, \lambda, \text{Chl}) \cdot \frac{b_b(\lambda, \text{Chl})}{a(\lambda, \text{Chl})}$$

$$\checkmark \quad E_d(0^-)(\theta_0, \lambda, \text{Chl})^* = E_d(0^+)(\theta_0, \lambda) \frac{1 - \bar{\rho}}{1 - \bar{r} R(\theta_0, \lambda, \text{Chl})^*}$$

$$\checkmark \quad Q(\theta_0, \theta', \Delta\phi, \lambda, \text{Chl}) = \frac{E_u(\theta_0, \lambda, \text{Chl})}{L_u(\theta_0, \theta', \Delta\phi, \lambda, \text{Chl})}$$

$$\checkmark \quad L_w(\theta_0, \theta_v, \Delta\phi, \lambda, \text{Chl}) = L_u(\theta_0, \theta', \Delta\phi, \lambda, \text{Chl}) \frac{1 - \rho_F(\theta')}{n^2}$$

$$L_w(\theta_0, \theta_v, \Delta\phi, \lambda, \text{Chl}) = E_d(0^+)(\theta_0, \lambda) \mathfrak{R}(\theta') \frac{f(\theta_0, \lambda, \text{Chl})}{Q(\theta_0, \theta', \Delta\phi, \lambda, \text{Chl})} \cdot \frac{b_b(\lambda, \text{Chl})}{a(\lambda, \text{Chl})}$$

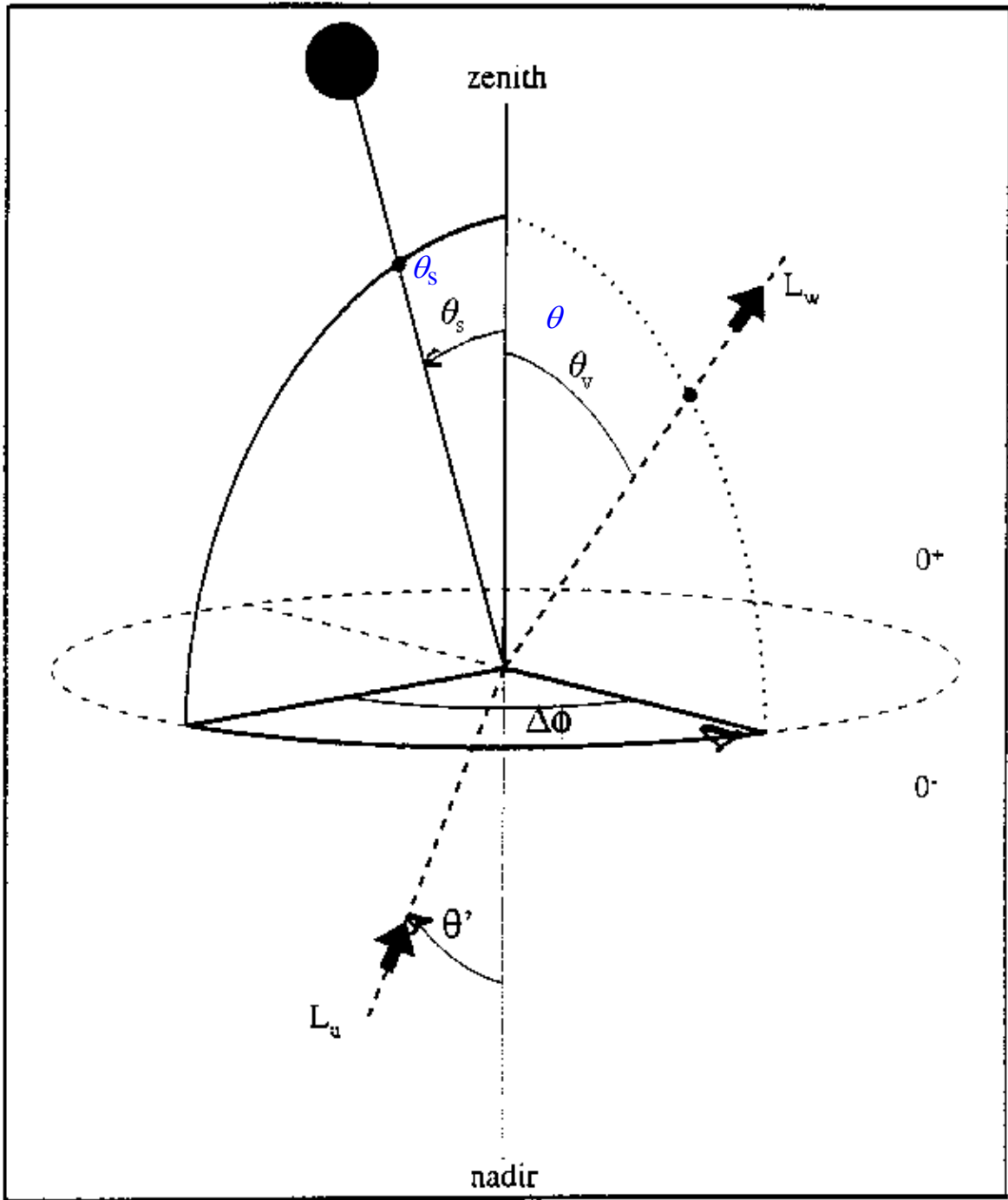
where $\sin(\theta_v) = n \cdot \sin(\theta')$

$$\text{and } \mathfrak{R}(\theta') = \left[\frac{(1 - \bar{\rho})}{(1 - \bar{r} R(\theta_0, \lambda, \text{Chl})^*)} \frac{(1 - \rho_F(\theta'))}{n^2} \right] \quad (\text{reflection} + \text{refraction})$$

Note that :

$$E_d(0^+)(\theta_0, \lambda) = \mu_0 \varepsilon_c F_0(\lambda) t(\theta_0, \lambda)$$

* *weak dependency on Chl*



L_u = in-water upwelling radiance

L_w = « water-leaving » radiance

$$L_w = L_u \frac{1 - \rho_F(\theta')}{n^2}$$

Useful expressions (λ and Chl omitted)

Slant water-leaving radiance :

$$L_w(\theta_0, \theta_v, \Delta\phi) = [\mu_0 \varepsilon_c F_0 t(\theta_0)] \mathfrak{R}(\theta') \frac{f(\theta_0)}{Q(\theta_0, \theta', \Delta\phi)} \cdot \frac{b_b}{a} \quad \text{Eq. (1)}$$

Vertically water-leaving radiance originating from nadir (subscript “n”):

(i.e., with $\theta_v = \theta' = \Delta\phi = 0$)

$$L_w(\theta_0, 0) = [\mu_0 \varepsilon_c F_0 t(\theta_0)] \mathfrak{R}_0 \frac{f(\theta_0)}{Q_n(\theta_0)} \cdot \frac{b_b}{a} \quad \text{Eq. (2)}$$

when, hypothetically, the sun zenith angle is naught $\theta_0 = 0$

$$L_w(0, 0) = [\mu_0 (= 1) \varepsilon_c F_0 t(0)] \mathfrak{R}_0 \frac{f_0}{Q_n(0)^*} \cdot \frac{b_b}{a} \quad \text{Eq. (3)}$$

* : also denoted Q_0

Gordon and Clark (1981) operational definition of satellite normalized water-leaving radiances

$$\text{Original definition : } [L_w]_N^{\text{GC}} = L_w^{\text{meas}} \frac{F_0}{E_d(0^+)} = \frac{L_w^{\text{meas}}}{\varepsilon_c \mu_0 t(\theta_0)}$$

➤ When using **Eq. (1)**

$$[L_w]_N^s = F_0 \mathfrak{R}(\theta') \frac{f(\theta_0)}{Q(\theta_0, \theta', \Delta\phi)} \cdot \frac{b_b}{a}$$

“s” for satellite

➤ When using **Eq. (3)**

$$[L_w]_N^{\text{ex}} = F_0 \mathfrak{R}_0 \frac{f_0}{Q_0} \cdot \frac{b_b}{a}$$

“ex” for exact (i.e., when $\theta' = 0$, $\theta_0 = 0$)

➤ Transforming $[L_w]_N^s$ into $[L_w]_N^{\text{ex}}$

$$[L_w]_N^{\text{ex}} = [L_w]_N^s \frac{\mathfrak{R}_0}{\mathfrak{R}(\theta')} \frac{f_0}{Q_0} \left(\frac{f(\theta_0)}{Q(\theta_0, \theta', \Delta\phi)} \right)^{-1} \quad \text{Eq. (4)}$$

From *in situ* determinations to “field” normalized water-leaving radiances

reminder :

Gordon and Clark (1981) operational definition

$$[L_w]_N^{GC} = L_w^{meas} \frac{F_0}{E_d(0^+)} = \frac{L_w^{meas}}{\varepsilon_c \mu_0 t(\theta_0)}$$

From **in-water** measurements (when measuring L_u at nadir $\rightarrow L_w^{inw}$)
(superscript “inw” for in-water, and “f” for field)

$$[L_w^{inw}]_N^f = \frac{L_w(\theta_0, \theta_v = 0)}{\mu_0 \varepsilon_c t(\theta_0)}$$

From **above-water** measurements (when measuring L_w at $\theta_v \neq 0$)
(superscript “abw” for above-water)

$$[L_w^{abw}]_N^f = \frac{L_w(\theta_0, \theta_v, \Delta\phi)}{\mu_0 \varepsilon_c t(\theta_0)}$$

➤ Transforming $[L_w^{abw}]_N^f$ into $[L_w^{inw}]_N^f$

$$[L_w^{inw}]_N^f = [L_w^{abw}]_N^f \frac{\mathfrak{R}_0}{\mathfrak{R}(\theta')} \frac{Q(\theta_0, \theta', \Delta\phi)}{Q_n(\theta_0)} \quad \text{Eq. (5)}$$

➤ Transforming $[L_w^{inw}]_N^f$ into $[L_w]_N^{ex}$

$$[L_w]_N^{ex} = [L_w^{inw}]_N^f \frac{f_0}{Q_n(0)} \left(\frac{f(\theta_0)}{Q_n(\theta_0)} \right)^{-1} \quad \text{Eq. (6)}$$

Summary/Conclusion

From satellites: $\ll nL_w \gg = [L_w]_N^s$

Still dependent on geometry,

- can be corrected for directional effects (Eq. 4), to provide $[L_w]_N^{ex}$ in a reversible way

From field measurements:

